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# Characterization of the defect density states in MoO<sub>x</sub> for c-Si solar cell applications

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## Abstract

Thin layers of MoO<sub>x</sub> have been deposited by thermal evaporation followed by post-deposition annealing. The density of states distributions of the MoO<sub>x</sub> films were extracted deconvoluting the absorption spectra, measured by a photothermal deflection spectroscopy setup, including the small polaron contribution. Results revealed a sub-band defect distribution centered 1.1 eV below the conduction band; the amplitude of this distribution was found to increase with post-deposition annealing temperature and film thickness.

## Keywords

*c-Si solar cell, photovoltaic, transition metal oxide, molybdenum oxide, density of states, small polaron.*

## 1. Introduction

Silicon-based heterojunction technology (HJT) is one of the most promising candidates for high performance and low-cost solar cells with world-record efficiency close to 27% in interdigitated back contact architecture [1]. The HJT exploits the excellent passivation properties of hydrogenated amorphous silicon (a-Si:H), although the use of doped a-Si:H has drawbacks such as parasitic absorption and low-thermal budget to cope with back-end metallization.

Replacing the p-type a-Si:H with a doping-free transition metal oxide (TMO) such as molybdenum oxide (MoO<sub>x</sub>), is a viable alternative allowing conversion efficiency up to 23.5% [2] rivaling the traditional contact despite its lower level of optimization. Moreover, the hole selectivity capability of the MoO<sub>x</sub> is exploited for other classes of electronic devices such as organic light-emitting diodes [3], organic photovoltaic cells [4], thin-film solar cells [5]. Thus, the optimization of this hole-selective layer has been investigated over the past years, highlighting the role of the defect density of states (DOS) [6–8]. Nonetheless, information on the DOS, linked to oxygen vacancies [9], lacks for TMOs.

We aim to fill this gap by providing insights into the MoO<sub>x</sub> defect density needed for accurate simulation and optimization of HJT solar cells with this TMO as a hole-selective contact. Therefore, chemical, morphological and optical characterizations were conducted on thin films of MoO<sub>x</sub>. Finally, the DOS of MoO<sub>x</sub> samples were extracted from the deconvolution of the absorption spectra [10,11], including the absorption related to excitation of electrons from small polaron states filled by electron transfer from the oxygen vacancies. A systematic study of the effects of both layers thickness and post-deposition annealing (PDA) treatments on the extracted DOS is furthermore presented.

## 2. Experimental

Thin films of molybdenum oxide were deposited by thermal evaporation on quartz substrates with desired thicknesses of 20, 50 and 100 nm. After deposition, the layers were annealed at different temperatures ( $T_{PDA}$  from 100 to 250 °C) for 30 minutes in ambient air. The desired thickness of the samples was confirmed by ex-situ ellipsometry and Rutherford backscattering spectrometry (RBS) which did not show, as displayed in Fig. 1, any significant variation between as-deposited and annealed samples.

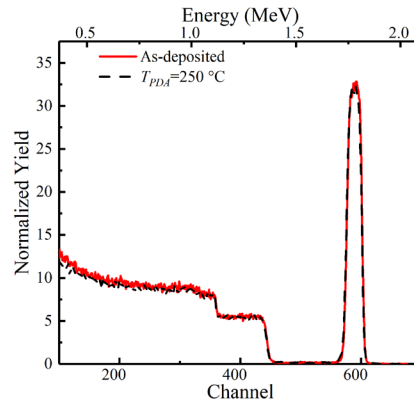


Fig. 1: RBS spectrum of the 20 nm  $\text{MoO}_x$  samples as-deposited and after annealing at 250°C for 30 minutes.

The structure and morphology of the samples were investigated through Raman spectroscopy and Scanning Electron Microscopy, respectively. The measurements show that all the samples were amorphous, independently on the PDA and the thickness, with no significant surface morphology variation. The chemical composition was studied by X-ray photoelectron spectroscopy (XPS). The optical properties were analyzed through ellipsometry and photothermal deflection spectroscopy (PDS).

### 3. Results

The Mo 3d binding region is presented in Fig. 2. From XPS measurement, the presence of a doublet structure with a spin-orbit separation of 3.0 eV is highlighted. The Mo 3d level is deconvoluted with two doublet components, a dominating one associated with the presence of  $\text{Mo}^{6+}$  [12] and a weak doublet shifted 2 eV to lower energy indicating the presence of  $\text{Mo}^{5+}$  (linked to the oxygen vacancies) [12].

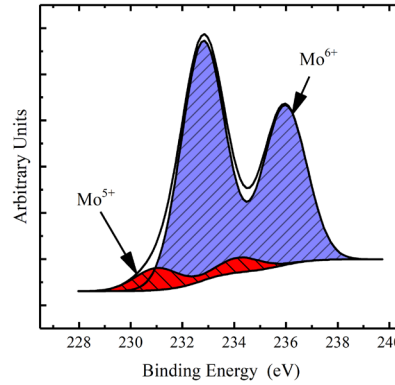


Fig.2: XPS spectrum of the Mo 3d binding region.

The absorption spectra, presented in Fig 3, decay exponentially from lower wavelength to a valley at about 490 nm and then rises to a peak around 900 nm. There are no significant differences between the curves in the wavelength range of 350-400 nm; while the peak observed in the Vis-NIR range tendentially grows as the annealing temperature rises. This increase in the absorption is referred to  $\text{MoO}_x$  reduction in the annealed samples. Interestingly, for the 20 nm thick  $\text{MoO}_x$  films (Fig. 3a), the spectra of the samples annealed at 200°C and 250 °C have a comparable magnitude, whilst, for the higher thicknesses, such behavior is less evident. Furthermore, for the 50 nm thick samples (Fig. 3b), the spectrum of the sample annealed at 100°C is similar to the as-deposited one; for the 100 nm case (Fig. 3c), such behavior is more evident since the as-deposited sample presents a spectrum higher than the sample annealed at 100°C.

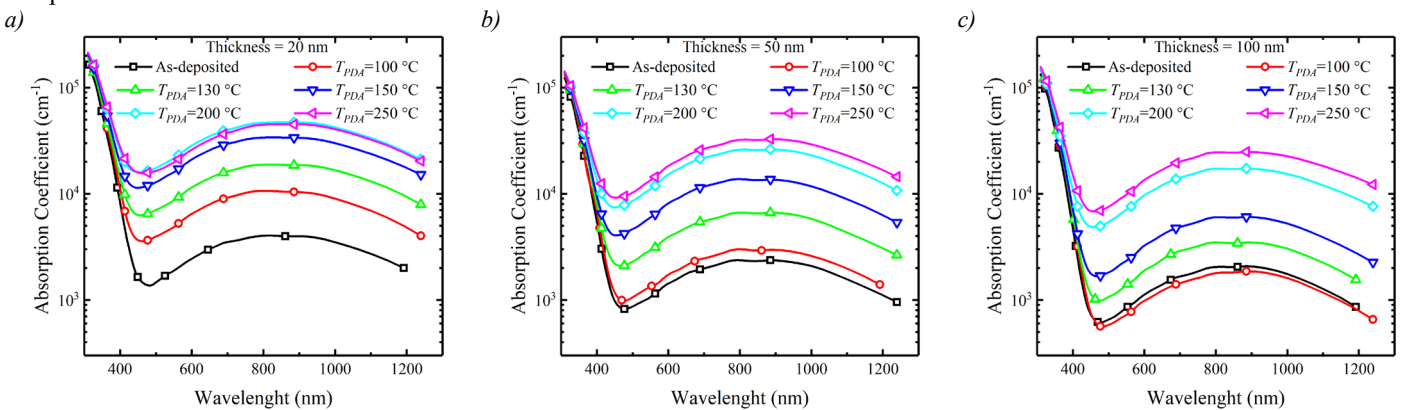


Fig.3: Absorption spectrum of a) 20 nm, b) 50 nm, c) 100 nm thick  $\text{MoO}_x$  films as deposited and after different PDA

#### 4. Model and discussion

The density of states of the MoO<sub>x</sub> was described as the superposition of different distributions, assuming parabolic distribution for the valence and conduction bands ( $N_{VB}$ ,  $N_{CB}$ ), exponential distribution for the valence and conduction band tails ( $N_{VBT}$ ,  $N_{CBT}$ ), and Gaussian distribution for the defects states in the bandgap ( $N_D$ ) [11]:

$$DOS(E) \rightarrow \begin{cases} N_{VB}(E) = N_V \sqrt{-E + E_G + E_{0V}} & E < 0 \\ N_{CB}(E) = N_C \sqrt{E - E_{0C}} & E \geq 0 \\ N_{VBT}(E) = N_V \sqrt{\frac{E_{0V}}{2}} \exp\left(-\frac{E + E_G}{E_{0V}}\right) & E > E_G \\ N_{CBT}(E) = N_C \sqrt{\frac{E_{0C}}{2}} \exp\left(\frac{E}{E_{0C}}\right) & E < 0 \\ N_D(E) = \frac{A_D}{\sqrt{2\pi W^2}} \exp\left[-\frac{1}{2} \left(\frac{E + E_D}{W}\right)^2\right] & E_G < E < 0 \end{cases} \quad (1)$$

with  $E$  energy (the zero-energy reference was considered at the edge of the conduction band),  $E_{0V}$  and  $E_{0C}$  valence and conduction band tails,  $E_G$  energy gap; for Gaussian defect distribution:  $A_D$  area,  $E_D$  mean position and  $W$  the FWHM. The densities of states of the valence ( $N_V$ ) and conduction ( $N_C$ ) band were  $7.92 \times 10^{17} \text{ cm}^{-3} \text{ eV}^{-3/2}$  and  $6.78 \times 10^{18} \text{ cm}^{-3} \text{ eV}^{-3/2}$ , respectively [11].

The peaks in the absorption spectra are attributed to the small polaron [13,14], a quasi-particle describing the interaction of a trapped electron with the surrounding atoms [14,15]. Furthermore, small polaron is a typical feature of TMOs [13,16–19].

Therefore, to achieve the best fit of the absorption spectra, we introduce an additive term in the set of equations (1) to consider the polaron contribution. The polaron absorption,  $\alpha_p$  function of the photon energy ( $h\nu$ ), is modeled with a weakly asymmetric Gaussian peak [11,13,18]:

$$\alpha_p(h\nu) = \frac{A_p}{h\nu} \exp\left(-\frac{(h\nu - 2E_p)^2}{8E_p E_{op}}\right) \quad (2)$$

with  $A_p$  polaron pre-exponential factor,  $E_p$  polaron binding energy, and  $E_{op}$  longitudinal-optical phonon energy. The values of the DOS distributions were extracted through the absorption spectra  $\alpha$  fitting through the one-electron approximation and including the polaron absorption  $\alpha_p$  [11]:

$$\alpha(h\nu) = \frac{C}{h\nu} \int N_i(E) F(E) N_f(E + h\nu) [1 - F(E + h\nu)] dE + \alpha_p(h\nu) \quad (3)$$

where  $N_i$  and  $N_f$  are the initial and final states,  $F$  is the Fermi-Dirac function. The constant  $C$  was calibrated at 4 eV resulting  $4.13 \cdot 10^{-31} \text{ cm}^5 \cdot \text{eV}^2$  [11]. The obtained DOS lays about 1.1 eV below the conduction band edge with an amplitude rising with both  $T_{PDA}$  and film thickness (Fig. 4a). Interestingly, the defect distribution amplitude rises as the thickness is reduced and for increasing  $T_{PDA}$ . A similar trend was found for the  $A_p$  (Fig. 4b), whilst the other parameters of the small polaron remained fixed.

The resulting coefficients are summarized in Table 1 and are limited for brevity to the 100 nm thick samples.

Table 1: Defect distributions and small polaron coefficients at different  $T_{PDA}$  for the 100 nm thick samples

$T_{PDA}$ (°C)	Thickness (nm)	$E_{0V}$ (meV)	$E_{0C}$ (meV)	$A_D$ ( $\text{eV}^{-1} \text{cm}^{-3}$ )	$E_D$ (eV)	$W$ (eV)	$A_p$ ( $\text{eV}^{-1} \text{cm}^{-3}$ )	$E_p$ (eV)	$E_{op}$ (meV)
As. dep.	100	100	141	$3.41 \times 10^{14}$	1.14	0.05	$2.46 \times 10^3$	0.80	52.5
100	100	117	191	$3.61 \times 10^{14}$	1.16	0.05	$2.10 \times 10^3$	0.78	43.9
130	100	115	269	$5.07 \times 10^{14}$	1.18	0.05	$4.39 \times 10^3$	0.81	52.2
150	100	118	243	$8.89 \times 10^{14}$	1.18	0.05	$7.67 \times 10^3$	0.80	52.5
200	100	128	182	$3.21 \times 10^{15}$	1.15	0.05	$2.01 \times 10^4$	0.79	55.8
250	100	83	101	$5.07 \times 10^{15}$	1.11	0.05	$2.82 \times 10^4$	0.77	54.6

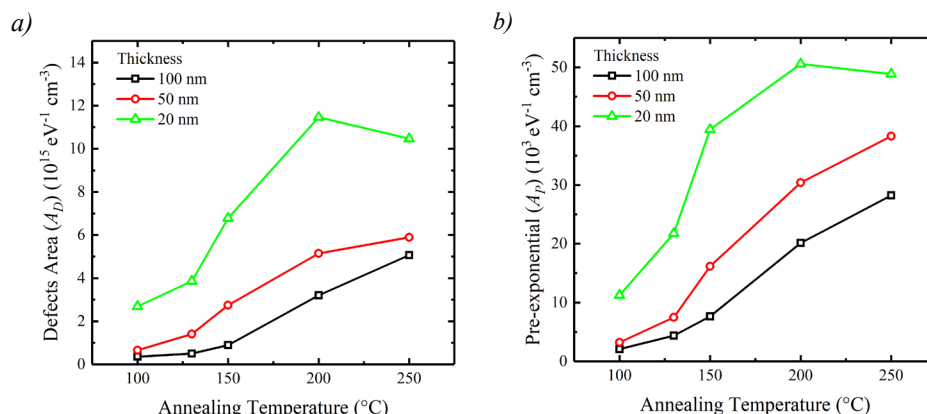


Fig. 4: Defects area (a) and polaron pre-exponential factor (b) against annealing temperature for different  $\text{MoO}_x$  films thicknesses

## 5. Conclusion

Thin films of  $\text{MoO}_x$  were synthesized by thermal evaporation, subsequently annealed, and then characterized to provide an insight into the DOS. The deconvolution of the absorption spectra resulted in a defect DOS distribution centered 1.1 eV below the conduction band edge with its amplitude increasing against both  $T_{PDA}$  and film thickness. The small polaron parameters, extrapolated from the optical measurements, revealed that both binding and longitudinal-optical phonon energy are independent of the thickness and the  $T_{PDA}$ , whilst the pre-exponential factor exhibits a similar trend to the amplitude of the defect distribution. The DOS characterization here employed for  $\text{MoO}_x$  films has shown to be a valuable method to determine the DOS and could be easily extended to other TMOs currently exploited as carrier selective contacts in solar cells.

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