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THE CONVERGING FACTOR FOR THE
MODIFIED BESSEL FUNCTION
OF THE SECOND KIND

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Naval Ship Research and Development Center
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DEPARTMENT OF THE NAVY
NAVAL SHIP RESEARCH AND DEVELOPMENT CENTER
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THE CONVERGING FACTOR FOR THE
MODIFIED BESSEL FUNCTION
OF THE SECOND KIND

by

John W. Wrench, Jr.

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ABSTRACT

The converging factor for a specific mathematical function, such as the modified Bessel function of the second kind considered in this report, is that factor by which the last term of a truncated series (usually asymptotic) approximating the function must be multiplied to compensate for the omitted terms. This converging factor for the aforementioned Bessel function is discussed herein in detail and is shown to be related to the corresponding factor for the probability integral. Tables of this factor and its reduced derivatives, correct to 30 decimal places, are included to expedite the application of this procedure to the evaluation of this Bessel function to high precision for arguments between 5 and 20, and specific examples of such applications are presented.

ADMINISTRATIVE INFORMATION

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INTRODUCTION

A variety of methods have been proposed in recent years for the calculation of Bessel functions as alternatives to the well-known use of power series for small arguments and asymptotic series for large arguments. These alternative procedures include recurrence relations, described by Abramowitz and Stegun¹ and by Goldstein and Thaler;² phase amplitude methods, also discussed by Goldstein and Thaler;³ quadrature methods, treated by Fettis,⁴ Luke,⁵ and Hunter;⁶ and continued fractions, discussed by Gargantini and Henrici.⁷

In addition to these methods, the use of converging factors to extend the precision attainable by asymptotic series for Bessel functions, in particular, has been advocated by several investigators including Airey⁸ and Dingle.⁹ Murnaghan¹⁰ and the writer¹¹ have investigated in detail the use of converging factors in the numerical evaluation to high precision of the probability integral and the exponential integral.

This report will show that the converging factor for the probability integral can be directly applied to the evaluation of the converging factor for the modified Bessel function of the second kind, $K_n(x)$.

As explicitly noted by Hunter,⁶ the evaluation of $K_n(x)$ by power series for even moderately large positive values of x presents special difficulty because of the loss of significant figures arising from the subtraction of nearly equal numbers. This computational difficulty

¹ References are listed on page 55.

can be avoided if an asymptotic series is used and the remainder resulting from truncating the series is closely estimated by use of a converging factor. Thus, for arguments x exceeding 5, such a procedure more than doubles the number of significant figures of $K_n(x)$ that can be attained by the conventional use of asymptotic series (that is, terminating the series at the least numerical term).

As emphasized by Gargantini and Henrici,⁷ the function K_n occupies a central position in the theory of Bessel functions, inasmuch as in the complex plane all other Bessel functions are expressible in terms of it. Accordingly, our attention in this study will be principally focussed upon the determination of the converging factor for this particular Bessel function.

A table of the basic converging factor $C_n(n)$ (or $\Lambda_n - \frac{1}{2}(n)$ in the notation of Dingle)

and its reduced derivatives rounded to 30 decimal places for $n = 10(1)40$ is included; this permits the calculation of $K_p(x)$ to a precision ranging from 15 decimal places when $x = 5$ to 42 decimal places when $x = 20$ and p is either 0 or 1.

THE ASYMPTOTIC SERIES FOR THE BESEL
FUNCTION $K_p(z)$ AND ITS CONVERGING FACTOR

The modified Bessel functions $I_p(z)$ and $K_p(z)$ satisfy the second-order linear differential equation

$$z^2 \frac{d^2 w}{dz^2} + z \frac{dw}{dz} - (z^2 + p^2) w = 0 \quad (1)$$

These functions may be distinguished according to their behavior when the argument z is large in absolute value; thus, the modified Bessel function of the first kind, $I_p(z)$, behaves asymptotically like $z^{-1/2} e^z$, whereas that of the second kind, $K_p(z)$, behaves asymptotically like $z^{-1/2} e^{-z}$.

If in Eq. (1) we make the change of variable

$$w(z) = x^p e^{-x/2} f(x), \quad x = 2z \quad (2)$$

we obtain after some simplification the differential equation

$$x \frac{d^2 f}{dx^2} + (2p + 1 - x) \frac{df}{dx} - (p + \frac{1}{2}) f = 0 \quad (3)$$

Since the confluent hypergeometric functions $f(a, c, x)$ satisfy Kummer's equation

$$x \frac{d^2 f}{dx^2} + (c - x) \frac{df}{dx} - af = 0 \quad (4)$$

we infer that

$$\begin{aligned} w(z) &= x^p e^{-x/2} f(p + \frac{1}{2}, 2p + 1, x) \\ &= (2z)^p e^{-z} f(p + \frac{1}{2}, 2p + 1, 2z) \end{aligned} \quad (5)$$

Now, if $\Psi(a, c, x)$ represents the solution of Eq. (4) which behaves like x^{-a} when x tends to infinity, we have the relation

$$K_p(z) = \pi^{1/2} (2z)^p e^{-z} \Psi(p + \frac{1}{2}, 2p+1, 2z) \quad (6)$$

expressing the modified Bessel function of the second kind of order p in terms of the confluent hypergeometric function Ψ .

To find the asymptotic series expansion of $\Psi(a, c, x)$ we set

$$\Psi(a, c, x) = x^{-a} \psi(a, c, x) \quad (7)$$

Differentiating this twice with respect to x and substituting the results in Eq. (4), we obtain

$$\frac{d^2\psi}{dx^2} - \left(1 + \frac{2a-c}{x}\right) \frac{d\psi}{dx} + \frac{a(a-c+1)}{x^2} \psi = 0 \quad (8)$$

Next, we assume a solution to Eq. (8) of the form

$$\psi = a_1 + \frac{a_2}{x} + \frac{a_3}{x^2} + \cdots + \frac{a_n}{x^{n-1}} + \cdots \quad (9)$$

If we differentiate this series twice with respect to x and substitute the results in Eq. (8), we find the coefficient of x^{-r} to be $(r-1)a_r + (r+a-2)(r+a-c-1)a_{r-1}$, which must vanish; hence, we infer that

$$\frac{a_r}{a_{r-1}} = - \frac{(r+a-2)(r+a-c-1)}{r-1} \quad (10)$$

Thus, the function $\psi(a, c, x)$ may be defined by the asymptotic series

$$\begin{aligned} \psi(a, c, x) &\sim 1 - \frac{a(a-c+1)}{1!x} + \frac{a(a+1)(a-c+1)(a-c+2)}{2!x^2} - \dots \\ &\sim \frac{1}{(a-1)!(a-c)!} \sum_{r=1}^{\infty} \frac{(r+a-2)!(r+a-c-1)!}{(r-1)!(-x)^{r-1}} \end{aligned} \quad (11)$$

We observe that if either a or $a-c+1$ is zero or a negative integer, this series terminates; indeed, we have

$$\begin{aligned}\Psi(-k, c, x) &= x^{-a} \psi(a, c, x) \\ &= (-1)^k k! L_k^{c-1}(x)\end{aligned}\quad (12)$$

where

$$L_k^n(x) = (k!)^{-1} x^{-n} e^x \left(\frac{d}{dx}\right)^k (x^{k+n} e^{-x}) \quad (13)$$

is the generalized Laguerre polynomial.

In order to derive an expression for the converging factor associated with the asymptotic series (11), we proceed as follows.

From the definition of the Beta function we have

$$\begin{aligned}B(z, w) &= \int_0^1 t^{z-1} (1-t)^{w-1} dt \\ &= \int_1^\infty (u-1)^{z-1} u^{-z-w} du\end{aligned}\quad (14)$$

where $u = (1-t)^{-1}$. Furthermore, the Beta function can be expressed in terms of the Gamma or the factorial function, as follows:

$$\begin{aligned}B(z, w) &= \frac{\Gamma(z) \Gamma(w)}{\Gamma(z+w)} \\ &= \frac{(z-1)!(w-1)!}{(z+w-1)!}\end{aligned}\quad (15)$$

Consequently, we have the relation

$$\frac{(w-1)!}{(z+w-1)!} = \frac{1}{(z-1)!} \int_1^\infty (u-1)^{z-1} u^{-z-w} du$$

or

$$\frac{(r+a-2)!}{(r-1)!} = \frac{1}{(-a)!} \int_1^\infty (u-1)^{-a} u^{-r} du \quad (16)$$

if we set $w = r+a-1$ and $z = -a+1$.

Therefore, the series in (11) can be written

$$\psi(a, c, x) \sim \frac{1}{(a-1)!(a-c)!(-a)!} \int_1^{\infty} (u-1)^{-a} u^{-1} du \sum_{r=1}^{\infty} \frac{(r+a-c-1)!}{(-ux)^{r-1}} \quad (17)$$

$$\sim \frac{1}{(a-1)!(a-c)!} \left\{ \sum_{r=1}^n \frac{(r+a-2)!(r+a-c-1)!}{(r-1)!(-x)^{r-1}} \right.$$

$$\left. + \frac{(n+a-c)!}{(-x)^n} \frac{1}{(-a)!} \int_1^{\infty} (u-1)^{-a} u^{-n-1} A_{n+a-c}(ux) du \right\} \quad (18)$$

where

$$A_{n+a-c}(ux) \sim 1 - \frac{n+a-c+1}{ux} + \frac{(n+a-c+1)(n+a-c+2)}{(ux)^2} - \dots \quad (19)$$

We observe that

$$\begin{aligned} A_{n+a-c}(ux) &= \frac{1}{(n+a-c)!} \left\{ (n+a-c)! - \frac{(n+a-c+1)!}{ux} + \frac{(n+a-c+2)!}{(ux)^2} - \dots \right\} \\ &= \frac{1}{(n+a-c)!} \int_0^{\infty} e^{-t} t^{n+a-c} \left(1 - \frac{t}{ux} + \frac{t^2}{(ux)^2} - \dots \right) dt \\ &= \frac{1}{(n+a-c)!} \int_0^{\infty} \frac{e^{-t} t^{n+a-c}}{1 + \frac{t}{ux}} dt \end{aligned} \quad (20)$$

which is identifiable with the first basic converging factor of Dingle, independently derived by Murnaghan and the present writer in a joint study of the exponential integral, wherein the notation $\Gamma_{n+a-c}(ux)$ was used for this converging factor.

If we write

$$\begin{aligned}\psi(a, c, x) = & \frac{1}{(a-1)!(a-c)!} \left\{ \sum_{r=1}^n \frac{(r+a-2)!(r+a-c-1)!}{(r-1)!(-x)^{r-1}} \right. \\ & \left. + \frac{(n+a-1)!(n+a-c)!}{n!(-x)^n} \Sigma_{n+a-c}(x) \right\} \quad (21)\end{aligned}$$

we infer from (18) that the converging factor for this series is given by

$$\Sigma_{n+a-c}(x) = \int_1^\infty (u-1)^{-a} u^{-n-1} \Lambda_{n+a-c}(ux) du / \int_1^\infty (u-1)^{-a} u^{-n-1} du \quad (22)$$

Since the factor u^{-n-1} in the integrand in (18) forms a rapidly decreasing sequence for increasing values of u , we expand the factor $\Lambda_{n+a-c}(ux)$ in ascending powers of $u-1$ and then integrate term by term. Thus, we write

$$\begin{aligned}\Lambda_{n+a-c}(ux) = & \Lambda_{n+a-c}(x) + x \Lambda_{n+a-c}^{(1)}(x) (u-1) + \dots \\ & + x^t \Lambda_{n+a-c}^{(t)}(x) (u-1)^t + \dots \quad (24)\end{aligned}$$

where

$$\overset{(k)}{\Lambda}_{n+a-c}(x) = \frac{1}{k!} \frac{d^k}{dx^k} \Lambda_{n+a-c}(ux) \Big|_{u=1} \quad (25)$$

which is the k^{th} reduced derivative of $\Lambda_{n+a-c}(ux)$, evaluated at $u = 1$.

Since

$$\begin{aligned}\int_1^\infty (u-1)^{-a} u^{-n-1} \Lambda_{n+a-c}^{(t)}(x) (u-1)^t du &= \Lambda_{n+a-c}^{(t)}(x) \int_1^\infty (u-1)^{t-a} u^{-n-1} du \\ &= \Lambda_{n+a-c}^{(t)}(x) \frac{(t-a)!(n-t+a-1)!}{n!} \quad (26)\end{aligned}$$

we can write

$$\begin{aligned}\psi(a, c, x) = & \frac{1}{(a-1)!(a-c)!} \left\{ \sum_{r=1}^n \frac{(r+a-2)!(r+a-c-1)!}{(r-1)!(-x)^{r-1}} \right. \\ & \left. + \frac{(n+a-1)!(n+a-c)!}{n!(-x)^n} \sum_{t=0}^{\infty} \frac{(t-a)!(n+a-1-t)!}{(-a)!(n+a-1)!} x^t \Lambda_{n+a-c}^{(t)}(x) \right\} \quad (27)\end{aligned}$$

which exhibits the converging factor $\Sigma_{n+a-c}(x)$ in Eq. (21) as an infinite series involving the basic converging factor $\Lambda_{n+a-c}(x)$ and its reduced derivatives.

Let us now consider again the modified Bessel function of the second kind. From Eqs. (6) and (7) we obtain

$$K_p(z) = \pi^{1/2} (2z)^{-1/2} e^{-z} \psi(p + \frac{1}{2}, 2p + 1, 2z) \quad (28)$$

where the principal square root is selected, in accordance with the stipulations $\arg w^{1/2} = \frac{1}{2} \arg w$ and $-\pi < \arg w \leq \pi$.

Thus, setting $a = p + \frac{1}{2}$, $c = 2p + 1$, and $x = 2z$ in (11), we deduce the well-known asymptotic expansion

$$\begin{aligned}K_p(x) \sim & \pi^{1/2} (2z)^{-1/2} e^{-z} \left\{ 1 + \frac{4p^2 - 1}{1!8z} + \frac{(4p^2 - 1)(4p^2 - 9)}{2!(8z)^2} + \dots \right. \\ & \left. + \frac{(4p^2 - 1)(4p^2 - 9) \dots [4p^2 - (2r-3)^2]}{(r-1)! (8z)^{r-1}} \dots \right\} \quad (29)\end{aligned}$$

This can be written in the form of the following terminating series:

$$\begin{aligned}
K_p(z) = & \pi^{1/2} (2z)^{-1/2} e^{-z} \left\{ \sum_{r=1}^{\lfloor p + \frac{1}{2} \rfloor} \frac{(r+p - \frac{3}{2})!}{(p + \frac{1}{2} - r)! (r-1)! (2z)^{r-1}} \right. \\
& + \frac{\cos \pi p}{\pi} \left[\sum_{r=\lfloor p + \frac{3}{2} \rfloor}^n \frac{(r+1 - \frac{3}{2})! (r-p - \frac{3}{2})!}{(r-1)! (-2z)^{r-1}} \right. \\
& \left. \left. + \frac{(n+p - \frac{1}{2})! (n-p - \frac{1}{2})!}{n! (-2z)^n} \Sigma_n \right] \right\} \quad (30)
\end{aligned}$$

where $\lfloor p + \frac{1}{2} \rfloor$ and $\lfloor p + \frac{3}{2} \rfloor$, respectively, represent the greatest integers not exceeding $p + \frac{1}{2}$ and $p + \frac{3}{2}$, and Σ_n designates the converging factor for this truncated series.

The coefficient $\frac{\cos \pi p}{p}$ arises from the relation

$$\begin{aligned}
(-p - \frac{1}{2})! (p - \frac{1}{2})! &= \Gamma(-p + \frac{1}{2}) \Gamma(p + \frac{1}{2}) \\
&= \frac{\pi}{\sin(p + \frac{1}{2})\pi} = \frac{\pi}{\cos \pi p} \quad (31)
\end{aligned}$$

From Eq. (27) we infer that the converging Σ_n can be evaluated by means of the series

$$\Sigma_n = \sum_{t=0}^{\infty} \frac{(t-p - \frac{1}{2})! (n+p - \frac{1}{2} - t)!}{(-p - \frac{1}{2})! (n+p - \frac{1}{2})!} (2z)^t A_{n-p-\frac{1}{2}}^{(t)} (2z) \quad (32)$$

Since

$$\begin{aligned}
 \frac{(t-p-\frac{1}{2})!}{(-p-\frac{1}{2})!} &= (t-p-\frac{1}{2})(t-p-\frac{3}{2}) \cdots (-p+\frac{1}{2}) \\
 &= (-1)^t (p-\frac{1}{2})(p-\frac{3}{2}) \cdots (p-t+\frac{3}{2})(p-t+\frac{1}{2}) \\
 &= (-1)^t \frac{(\frac{p}{2}-\frac{1}{2})!}{(\frac{p}{2}-t)!}
 \end{aligned} \tag{33}$$

we can write alternatively

$$\Sigma_n = \sum_{t=0}^{\infty} \frac{(\frac{p}{2}-\frac{1}{2})!(n+p-\frac{1}{2}-t)!}{(\frac{p}{2}-t)!(n+p-\frac{1}{2})!} (-2z)^t A_{n-p-\frac{1}{2}}^{(t)}(2z) \tag{34}$$

If in Eq. (20) we make the substitutions $a = p + \frac{1}{2}$

$c = 2p+1$, and $ux = 2z$, we obtain for the basic converging factor

$A_{n-p-\frac{1}{2}}$ the expression

$$A_{n-p-\frac{1}{2}}(2z) = \frac{1}{(n-p-\frac{1}{2})!} \int_0^{\infty} \frac{e^{-t} t^{n-p-\frac{1}{2}}}{1 + \frac{t}{2z}} dt = C_{n-p}(2z) \tag{35}$$

where $C_n(y)$ designates the converging factor for the probability integral, as developed by Murnaghan.

To derive Eq. (35) we proceed as follows: If $\text{erfc}(x)$ denotes the complementary error function, we have by definition

$$\begin{aligned}
 \text{erfc}(x) &= \int_x^{\infty} e^{-t^2} dt \\
 &= \frac{1}{2} \int_y^{\infty} e^{-u} u^{-1/2} du ,
 \end{aligned} \tag{36}$$

where $u = t^2$ and $y = x^2$.

Furthermore, if we set $u = v + y$, we obtain

$$\begin{aligned} \operatorname{erfc}(x) &= \frac{e^{-y}}{2y^{1/2}} \int_0^{\infty} e^{-v} \left(1 + \frac{v}{y}\right)^{-1/2} dv \\ &= \frac{e^{-y}}{2y^{1/2}} C(y), \text{ say.} \end{aligned} \quad (37)$$

Then repeated integration by parts yields

$$\begin{aligned} C(y) &= 1 - \frac{1}{2y} + \frac{1 \cdot 3}{(2y)^2} - \dots + (-1)^{n-1} \frac{1 \cdot 3 \cdots (2n-3)}{(2y)^{n-1}} \\ &\quad + (-1)^n \frac{1 \cdot 3 \cdots (2n-1)}{(2y)^n} C_n(y) \end{aligned} \quad (38)$$

where the converging factor $C_n(y)$ is given by

$$C_n(y) = \int_x^{\infty} e^{-v} \left(1 + \frac{v}{y}\right)^{-n - \frac{1}{2}} dv \quad (39)$$

Moreover, from the definition of the exponential integral we have

$$\begin{aligned} -\operatorname{Ei}(-x) &= \int_x^{\infty} e^{-t} t^{-1} dt \\ &= \frac{e^{-x}}{x} \int_0^{\infty} e^{-u} \left(1 + \frac{u}{x}\right)^{-1} du \end{aligned} \quad (40)$$

where $x = t-u$.

Thus, if we set $-\operatorname{Ei}(-x) = \frac{e^{-x}}{x} A(x)$, then repeated integration by parts yields

$$\begin{aligned} A(x) &= 1 - \frac{1}{x} + \frac{\frac{2!}{2}}{x^2} - \cdots + (-1)^{n-1} \frac{\frac{(n-1)!}{n-1}}{x^{n-1}} \\ &\quad + (-1)^n \frac{n!}{x^n} A_n(x) \end{aligned} \quad (41)$$

where the converging factor $A_n(x)$ is given by

$$A_n(x) = \int_0^\infty e^{-u} \left(1 + \frac{u}{x}\right)^{-n-1} du \quad (42)$$

Comparison of Eqs. (39) and (42) reveals that

$$A_{n-\frac{1}{2}}(x) = C_n(x) \quad (43)$$

On the other hand, if we write

$$\left(1 + \frac{u}{x}\right)^{-1} = 1 - \frac{u}{x} + \cdots + (-1)^{n-1} \frac{u^{n-1}}{x^{n-1}} + (-1)^n \frac{u^n}{1 + \frac{u}{x}} \quad (44)$$

introduce this finite series in the integrand of Eq. (40), and integrate term by term, we obtain the series in Eq. (41), where $A_n(x)$ now assumes the equivalent form

$$A_n(x) = \frac{1}{n!} \int_0^\infty \frac{e^{-u} u^n}{1 + \frac{u}{x}} du \quad (45)$$

and thus the validity of Eq. (35) is finally established.

**THE DIFFERENCE AND DIFFERENTIAL EQUATIONS
SATISFIED BY THE CONVERGING FACTOR $C_n(y)$**

From Eq. (38) we infer that $C_n(y)$ satisfies the difference equation

$$C_n(y) = 1 - \frac{2n+1}{2y} C_{n+1}(y) \quad (46)$$

as noted by Murnaghan¹⁰.

To derive the differential equation satisfied by $C_n(y)$ we differentiate both sides of Eq. (39) with respect to y . This yields

$$\begin{aligned} \frac{d}{dy} C_n(y) &= \frac{m}{y^2} \int_0^\infty e^{-v} v \left(1 + \frac{v}{y}\right)^{-n-\frac{3}{2}} dv \\ &= \frac{m}{y} \{ C_n(y) - C_{n+1}(y) \} \\ &= \left(1 + \frac{m}{y}\right) C_n(y) - 1 \end{aligned} \quad (47)$$

by virtue of Eq. (46). Here we have replaced $(2n+1)/2$ by m , for typographical simplicity.

Then from Eq. (47) we obtain by differentiation the relation

$$\begin{aligned} y \frac{d^2}{dy^2} C_n(y) &= (y+m) \frac{d}{dy} C_n(y) - \frac{m}{y} C_n(y) \\ &= (y+m-1) \frac{d}{dy} C_n(y) + C_n(y) - 1 . \end{aligned} \quad (48)$$

Continuing this process, we find

$$y \frac{d^3}{dy^3} C_n(y) = (y+m-2) \frac{d^2}{dy^2} C_n(y) + 2 \frac{d}{dy} C_n(y) \quad (49)$$

and in general

$$y \frac{d^k}{dy^k} C_n(y) = (y+m-k+1) \frac{d^{k-1}}{dy^{k-1}} C_n(y) + (k-1) \frac{d^{k-2}}{dy^{k-2}} C_n(y) \quad (50)$$

if $k \geq 3$.

These derivatives are required in the Taylor series expansion

$$C_n(y_0+h) = C_n(y_0) + d_1 h + d_2 h^2 + \dots \quad (51)$$

where

$$d_j = [\frac{1}{j!} \frac{d^j}{dy^j} C_n(y)]_{y=y_0} \quad (52)$$

is the j^{th} reduced derivative of $C_n(y)$, evaluated at $y = y_0$.

From Eqs. (47)-(50) we then obtain the following equivalent recurrence relations among the reduced derivatives:

$$y_0 d_1 = (y_0+m)d_0 - y_0 \quad (53)$$

$$2y_0 d_2 = (y_0+m-1)d_1 + d_0 - 1 \quad (54)$$

$$3y_0 d_3 = (y_0+m-2)d_2 + d_1 \quad (55)$$

.

$$ky_0 d_k = (y_0+m-k+1)d_{k-1} + d_{k-2} \quad (k \geq 3) \quad (56)$$

Thus we have derived a systematic procedure for finding the successive reduced derivatives of the converging factor $C_n(y)$ (or $\Lambda_{n-\frac{1}{2}}(y)$ in the notation of Dingle) in terms of that factor.

In using Eq. (30) to calculate $K_p(z)$, it is generally most convenient to select n so that the asymptotic series is truncated at its least term numerically. Hence, if we are given the order p of the Bessel function and the argument z , we determine n such that

$$(n+p - \frac{1}{2})(n-p - \frac{1}{2}) \leq 2nz \quad (57)$$

which implies

$$n \leq z + \frac{1}{2} + (z^2 + z + p^2)^{1/2} \quad (58)$$

When $p = \pm \frac{1}{2}$, this reduces to the simpler inequality

$n \leq 2z + 1$; consequently, when $p = 0$ or 1 , we can select n to be the greatest integer not exceeding $2z+1$ if we assume z to be real and positive.

In that case, the appropriate converging factor is $C_n(n)$, which has been tabulated by Murnaghan¹⁰ to 63 decimal places for integer values of n from 2 to 64, inclusive.

These fundamental data have been used to calculate correct to 30 decimal places the reduced derivatives d_j for $y_0 = n = 10(1)40$; that is, for successive integer values of the argument from 10 to 40, inclusive. These values are tabulated in the Appendix.

THE ASYMPTOTIC SERIES FOR THE CONVERGING

FACTOR $C_n(n)$

In his study of the probability integral, Murnaghan¹⁰ derived the Airey asymptotic series for the converging factor $C_n(n + \frac{1}{2} + h)$, and then by setting $h = -\frac{1}{2}$, he deduced therefrom the corresponding asymptotic series for the converging factor $C_n(n)$.

This result can be written

$$2C_n(n) \sim 1 + \sum_{i=1}^{\infty} \frac{c_i}{(4n+2)^i} \quad (59)$$

where for convenient reference we list here in Table 1 the exact values of the first 30 coefficients c_i , which have been taken from the more extended Table 2 on page 38 of Murnaghan's report¹⁰

It should be pointed out here that the first 25 terms suffice to yield an approximation to $C_n(n)$ that is correct to more than 30 decimal places when $n = 40$. Thus by means of series (59) and the relations (53)-(56), one can readily extend the range of the tables in the Appendix, in order to accommodate values of n exceeding 40, if such are required.

TABLE 1

Coefficients in the Asymptotic Series $2C_n(n) \sim 1 + \frac{c_1}{4m} + \frac{c_2}{(4m)^2} + \frac{c_3}{(4m)^3} + \dots$

i	c_i
1	0
2	-1
3	-1
4	4
5	-21
6	-23
7	916
8	-6619
9	-3099
10	6 40760
11	-72 98875
12	71 97679
13	10988 76024
14	-1 83598 69769
15	5 79797 07895
16	370 89637 19852
17	-8723 83728 95349
18	52107 67357 60217
19	21 41277 71661 78716
20	-696 02236 33844 34419
21	6549 63005 10513 05805
22	1 91213 38271 60645 86192
23	-85 96151 42501 57889 82715
24	1159 74216 37624 11668 68319
25	24305 68772 33650 41843 64656
26	-15 47926 19401 73625 04069 29169
27	282 97994 75909 84357 78487 50447
28	4037 22733 27480 09248 96541 72372
29 -3	88039 12428 77823 20544 49772 83413
30 92	71310 71991 43807 59763 38256 22729

APPLICATIONS

As the first illustration of the use of the converging factor in evaluating $K_n(x)$ we take $n = 0$ and $x = 2\pi$, which is a relatively small argument for the effective utilization of the conventionally truncated asymptotic series. The value of $K_0(2\pi)$ to nine decimal places has been included in a table published by Olver¹² for use in aerodynamic calculations of interference on lifting surfaces in rectangular wind tunnels.

We evaluate $K_0(2\pi)$ by means of the series

$$K_0(2\pi) = \frac{e^{-2\pi}}{2} \left\{ 1 - \frac{1}{1!16\pi} + \frac{(3!!)^2}{2!(16\pi)^2} - \frac{(5!!)^2}{3!(16\pi)^3} \right. \\ \left. + \dots + \frac{(23!!)^2}{12!(16\pi)^{12}} - \frac{(25!!)^2}{13!(16\pi)^{13}} \Sigma_{13} \right\}$$

where the symbol $(2k-1)!!$ represents the product $1 \cdot 3 \cdots (2k-1)$, and the converging factor Σ_{13} can be calculated from the series

$$\Sigma_{13} = C_{13}(4\pi) + \frac{4\pi}{25} C_{13}^{(1)}(4\pi) + \frac{1 \cdot 3}{23 \cdot 25} (4\pi)^2 C_{13}^{(2)}(4\pi) + \dots$$

We find the value of $C_{13}(4\pi)$ from the series

$$C_{13}(4\pi) = C_{13}(13) + d_1 h + d_2 h^2 + \dots$$

where

$$h = -13 + 4\pi = -0.43362\ 93856\ 40827\ 046$$

and $C_{13}(13)$, d_1 , d_2 , ... (for $n = 13$) are tabulated in the Appendix.

We thus find the approximation

$$C_{13}(4\pi) = 0.49150\ 20002\ 93166\ 9\dots,$$

and then by means of formulas (47)-(50) we deduce the values

$$C_{13}^{(1)}(4\pi) = 0.01952\ 05672\ 75096$$

$$C_{13}^{(2)}(4\pi) = -0.376\ 34752\ 45781$$

$$C_{13}^{(3)}(4\pi) = .0304115\ 66724$$

$$C_{13}^{(4)}(4\pi) = -.0512332\ 67319$$

$$C_{13}^{(5)}(4\pi) = -.07508\ 95093$$

$$C_{13}^{(6)}(4\pi) = .082136555$$

$$C_{13}^{(7)}(4\pi) = .0_{10}91198$$

To seven decimal places, we infer that $\Sigma_{13} = 0.50074\ 87$
and consequently

$$\frac{(25!)^2}{13!(16\pi)^{13}} \Sigma_{13} = 0.063843\ 518\dots$$

Combining this with the earlier terms in the asymptotic series, we obtain the result

$$2e^{2\pi} K_0(2\pi) = 0.98164\ 65536\ 976\dots$$

whence

$$K_0(2\pi) = 0.00091\ 65843\ 60904\ 39\dots$$

Use of the power series for $K_0(x)$ leads to the series

$$K_0(2\pi) = -(\ln \pi + \gamma) I_0(2\pi) + \frac{\pi^2}{(1!)^2} + (1 + \frac{1}{2}) \frac{\pi^4}{(2!)^2} + (1 + \frac{1}{2} + \frac{1}{3}) \frac{\pi^6}{(3!)^2} + \dots$$

where $\gamma = 0.57721\ 56649\ 01532\ 86060\dots$ is Euler's constant and

$$I_0(2\pi) = 1 + \frac{\pi^2}{(1!)^2} + \frac{\pi^4}{(2!)^2} + \frac{\pi^6}{(3!)^2} + \dots$$

If we evaluate 22 terms of each series to 21 decimal places, we obtain

$$I_0(2\pi) = 87.10851\ 06533\ 90810\ 99853\dots \text{ and}$$

$$K_0(2\pi) = 0.00091\ 65843\ 60904\ 37031\dots$$

correct to 20 decimal places.

Consequently, the value of $K_0(2\pi)$ found by means of the converging factor is too large by less than 2×10^{-17} . This accuracy represents a gain of seven decimal places beyond that obtainable by the standard use of the asymptotic series in this case.

As a second example of the effectiveness of the converging-factor method in the calculation of $K_n(x)$ to high precision, we evaluate $K_1(10)$ to more than 25 decimal places by that procedure.

If we truncate the appropriate asymptotic series at the least numerical term and introduce the corresponding converging factor, we obtain

$$\begin{aligned} K_1(10) &= \left(\frac{\pi}{20}\right)^{1/2} e^{-10} \left\{ 1 + \frac{3!!}{80} - \frac{5!!}{2!} \frac{1}{80^2} \right. \\ &\quad + \frac{3!!7!!}{3!} \frac{1}{80^3} - \frac{5!!9!!}{4!} \frac{1}{80^4} + \dots \\ &\quad \left. + \frac{39!!43!!}{21!} \frac{1}{80^{21}} \Sigma_{20} \right\} \end{aligned}$$

where the converging factor Σ_{20} is computed from the series

$$\begin{aligned}\Sigma_{20} = & C_{20}(20) - \frac{20}{43} C_{20}^{(1)}(20) - \frac{20^2}{41 \cdot 43} C_{20}^{(2)}(20) \\ & - \frac{3 \cdot 20^3}{39 \cdot 41 \cdot 43} C_{20}^{(3)}(20) - \frac{3 \cdot 5 \cdot 20^4}{37 \cdot 39 \cdot 41 \cdot 43} C_{20}^{(4)}(20) - \dots\end{aligned}$$

Using the values of $C_{20}(20)$ and the reduced derivatives $C_{20}^{(k)}(20)$ ($= d_k$) tabulated in the Appendix, we calculate from the last series the approximation

$$\Sigma_{20} = 0.49424 \ 91283 \ 39$$

Then we find

$$\frac{39!!43!!}{21!} \ \frac{1}{80^{21}} \ \Sigma_{20} = 0.09^{189149} \ 702220 \dots$$

and combining this with the sum of the first 21 terms of the series within the braces, we obtain

$$\left(\frac{20}{\pi}\right)^{1/2} e^{10} K_1(10) = 1.03641 \ 84932 \ 28924 \ 58809 \ 9\dots$$

whence

$$K_1(10) = 0.00001 \ 86487 \ 73453 \ 82558 \ 45968 \ 10\dots$$

This approximation to the value of $K_1(10)$ was checked by means of power series, which entailed the calculation of $I_1(10)$ also.

Thus, evaluating 35 terms of the series

$$I_1(10) = 5 + \frac{5^3}{1!2!} + \frac{5^5}{2!3!} + \frac{5^7}{3!4!} + \dots$$

we find

$$I_1(10) = 2670.98830 \ 37012 \ 54654 \ 34103 \ 19667 \ 72152 \ 5\dots$$

and then substituting this value in the series

$$K_1(10) = (\ln 5 + \gamma) I_1(10) + \frac{1}{10} - \frac{1}{2} \left\{ \frac{5}{0!1!} + \frac{5^3}{1!2!} (1+1+\frac{1}{2}) \right.$$
$$\left. + \frac{5^5}{2!3!} (1+\frac{1}{2}+1+\frac{1}{2}+\frac{1}{3}) + \dots \right\}$$

we obtain

$$K_1(10) = 0.00001 \ 86487 \ 73453 \ 82558 \ 45968 \ 168581 \dots$$

provided we evaluate 36 terms of the series within the braces, and approximate each term to at least 32 decimal places.

Comparison of this latter value with the earlier one reveals that the use of the converging factor here suffices to yield the value of $K_1(10)$ to within 7×10^{-27} . This is a gain in accuracy of 12 decimal places over that obtainable by use of the asymptotic series without the converging factor.

APPENDIX

**TABLE OF THE CONVERGING FACTOR $C_n(n)$ ($=d_0$)
AND ITS REDUCED DERIVATIVES, d_i ,
TO 30 DECIMAL PLACES FOR $n = 10(1)40$**

n = 10

<i>d</i> ₀	0.49971	03665	11039	76983	07766	52551
<i>d</i> ₁	.02440	62513	47631	52815	30921	37729
<i>d</i> ₂	-.0 ₂ 121	83866	10507	27155	91963	33087
<i>d</i> ₃	.0 ₄ 6	22033	01774	90014	35986	65854
<i>d</i> ₄	-.0 ₅	32457	20736	16297	61554	92016
<i>d</i> ₅		.0 ₆ 1729	78192	56422	07406	60952
<i>d</i> ₆		-.0 ₈ 94	09312	52362	59112	54121
<i>d</i> ₇		.0 ₉ 5	22045	15673	77861	06803
<i>d</i> ₈		-.0 ₁₀	29521	28634	58224	85154
<i>d</i> ₉			.0 ₁₁ 1700	32308	23889	44915
<i>d</i> ₁₀			-.0 ₁₃ 99	67570	89834	96186
<i>d</i> ₁₁			.0 ₁₄ 5	94298	30732	93045
<i>d</i> ₁₂			-.0 ₁₅	36014	47482	26769
<i>d</i> ₁₃				.0 ₁₆ 2216	73285	64350
<i>d</i> ₁₄				-.0 ₁₇ 138	49270	28530
<i>d</i> ₁₅				.0 ₁₉ 8	77686	85859
<i>d</i> ₁₆				-.0 ₂₀	56387	45352
<i>d</i> ₁₇					.0 ₂₁ 3670	25481
<i>d</i> ₁₈					-.0 ₂₂ 241	89756
<i>d</i> ₁₉					.0 ₂₃ 16	13427
<i>d</i> ₂₀					-.0 ₂₄ 1	08848
<i>d</i> ₂₁						.0 ₂₆ 7424
<i>d</i> ₂₂						-.0 ₂₇ 512
<i>d</i> ₂₃						.0 ₂₈ 36
<i>d</i> ₂₄						-.0 ₂₉ 3

n = 11

d_0	0.49975	89636	36729	31597	99366	68663
d_1	.02223	42438	02400	87359	53250	04083
d_2	-0.2100	93088	45984	17735	09352	61071
d_3	.044	67700	74537	12842	12470	34913
d_4	-0.5	22111	90738	86757	12867	74552
d_5	.061066	00833	96506	09534	85558	
d_6	-0.852	37517	34059	09787	99656	
d_7	.092	62101	27071	83091	33652	
d_8	-0.10	13351	67777	79225	82137	
d_9	.012691	93881	75599	16088		
d_{10}	-0.1336	45912	49169	42863		
d_{11}	.0141	95206	40999	84548		
d_{12}	-0.15	10613	93012	66065		
d_{13}	.017585	73527	04132			
d_{14}	-0.1832	78860	42707			
d_{15}	.0191	86080	08128			
d_{16}	-0.20	10700	33987			
d_{17}	.022623	14370				
d_{18}	-0.2336	73257				
d_{19}	.0242	19066				
d_{20}	-0.25	13211				
d_{21}	.027805					
d_{22}	-0.2850					
d_{23}	.0293					

n = 12

<i>d</i> ₀	0.49979	62861	67303	14262	69706	94482
<i>d</i> ₁	.02041	74175	91577	24953	00651	67902
<i>d</i> ₂	-0.384	97666	84609	64555	90207	44159
<i>d</i> ₃	.043	60463	10773	89512	36694	00676
<i>d</i> ₄	-0.5	15577	29228	56042	50026	79784
<i>d</i> ₅		.07685	47693	14010	68519	07752
<i>d</i> ₆		-0.830	70127	94900	47415	34425
<i>d</i> ₇		.091	39884	83432	76087	32392
<i>d</i> ₈			-0.116480	65987	78290	49141
<i>d</i> ₉			.012305	12913	28095	31681
<i>d</i> ₁₀			-0.1314	59298	59940	10901
<i>d</i> ₁₁			.015	70856	69386	08891
<i>d</i> ₁₂				-0.163491	20300	19381
<i>d</i> ₁₃				.017174	46574	57478
<i>d</i> ₁₄				-0.198	83837	45586
<i>d</i> ₁₅				.020	45368	22938
<i>d</i> ₁₆					-0.212358	53790
<i>d</i> ₁₇					.022124	12087
<i>d</i> ₁₈					-0.246	60940
<i>d</i> ₁₉					.025	35596
<i>d</i> ₂₀						-0.261938
<i>d</i> ₂₁						.027107
<i>d</i> ₂₂						-0.296

n = 13

<i>d</i> ₀	0.49982	55690	82240	27535	66822	29194
<i>d</i> ₁	.01887	51985	13797	48438	09291	59510
<i>d</i> ₂	-.0 ₃ 72	52641	85227	84126	65240	07819
<i>d</i> ₃	.0 ₄ 2	83647	68608	08649	61818	19691
<i>d</i> ₄	-.0 ₅	11286	94671	88093	47356	00867
<i>d</i> ₅		.0 ₇ 456	79053	70408	40712	43080
<i>d</i> ₆		-.0 ₈ 18	79423	29798	88103	06085
<i>d</i> ₇		.0 ₁₀	78581	05599	24556	04048
<i>d</i> ₈		-.0 ₁₁ 3337		42986	66915	96415
<i>d</i> ₉		.0 ₁₂ 143		91968	76808	63849
<i>d</i> ₁₀		-.0 ₁₄ 6		29873	33252	03685
<i>d</i> ₁₁		.0 ₁₅		27965	44602	44777
<i>d</i> ₁₂			-.0 ₁₆ 1259		03153	29549
<i>d</i> ₁₃			.0 ₁₈ 57		45259	64298
<i>d</i> ₁₄			-.0 ₁₉ 2		65616	19844
<i>d</i> ₁₅			.0 ₂₀		12436	19058
<i>d</i> ₁₆				-.0 ₂₂ 589		42311
<i>d</i> ₁₇				.0 ₂₃ 28		26809
<i>d</i> ₁₈				-.0 ₂₄ 1		37127
<i>d</i> ₁₉					.0 ₂₆ 6726	
<i>d</i> ₂₀					-.0 ₂₇ 333	
<i>d</i> ₂₁					-.0 ₂₈ 17	
<i>d</i> ₂₂					-.0 ₂₉ 1	

n = 14

<i>d</i> ₀	0.49984	89654	04695	36092	61517	24574
<i>d</i> ₁	.01754	96795	73844	12759	96660	10739
<i>d</i> ₂	-0.362	62445	11235	39750	29654	63575
<i>d</i> ₃	.0.42	27190	48240	62128	02638	38714
<i>d</i> ₄		-0.68376	56805	34919	38613	85292
<i>d</i> ₅		.0.7313	77950	13665	75808	55701
<i>d</i> ₆		-0.811	93749	72783	02465	62813
<i>d</i> ₇		.0.10	46107	97204	56840	12167
<i>d</i> ₈			-0.111807	39579	32860	74118
<i>d</i> ₉			.0.1371	87585	93914	24546
<i>d</i> ₁₀			-0.142	89868	95368	09253
<i>d</i> ₁₁			.0.15	11850	71620	80866
<i>d</i> ₁₂				-0.17490	96083	35679
<i>d</i> ₁₃				.0.1820	60363	98583
<i>d</i> ₁₄				-0.20	87553	27335
<i>d</i> ₁₅					.0.213765	91201
<i>d</i> ₁₆					-0.22163	89938
<i>d</i> ₁₇					.0.247	21500
<i>d</i> ₁₈					-0.25	32114
<i>d</i> ₁₉						.0.261445
<i>d</i> ₂₀						-0.2866
<i>d</i> ₂₁						.0.293

n = 15

d_0	0.49986	79530	13128	16639	23752	27736
d_1	.01639	81711	26693	93833	11629	63063
d_2	-0 ₃ 54	61999	58313	35509	46105	78730
d_3	.0 ₄ 1	84771	62550	29595	85502	54872
d_4		-0 ₆ 6346	33136	33693	72413	09496
d_5		.0 ₇ 221	25125	83156	16220	74043
d_6		-0 ₉ 7	82693	64035	68430	93571
d_7		-0 ₁₀	28086	96802	68472	97920
d_8		-0 ₁₁ 1022	08243	10494	29937	
d_9		.0 ₁₃ 37	70454	31721	12032	
d_{10}		-0 ₁₄ 1	40956	50189	93475	
d_{11}			.0 ₁₆ 5338	46077	74217	
d_{12}			-0 ₁₇ 204	75842	63312	
d_{13}			.0 ₁₉ 7	95092	25143	
d_{14}			-0 ₂₀	31246	32492	
d_{15}				.0 ₂₁ 1242	34618	
d_{16}				-0 ₂₃ 49	95816	
d_{17}				.0 ₂₄ 2	03119	
d_{18}					-0 ₂₆ 8347	
d_{19}					.0 ₂₇ 347	
d_{20}					-0 ₂₈ 15	
d_{21}					.0 ₂₉ 1	

n = 16

<i>d</i> ₀	0.49988	35735	94367	86390	45222	58666
<i>d</i> ₁	.01538	85088	63559	72355	60608	37916
<i>d</i> ₂	-0.348	05749	12609	40137	74862	92093
<i>d</i> ₃	.041	52286	25603	60378	21401	86022
<i>d</i> ₄		-0.64895	38395	36077	81929	81319
<i>d</i> ₅		.07159	59766	69777	00455	02730
<i>d</i> ₆		-0.95	27550	11637	60618	92253
<i>d</i> ₇		.010	17675	79119	46821	90697
<i>d</i> ₈		-0.12600	13625	71223	90855	
<i>d</i> ₉		.01320	64203	39943	30658	
<i>d</i> ₁₀		-0.15	71905	28640	97627	
<i>d</i> ₁₁			.0162535	99122	28034	
<i>d</i> ₁₂			-0.1890	52851	62473	
<i>d</i> ₁₃			.0193	26998	38449	
<i>d</i> ₁₄			-0.20	11948	13896	
<i>d</i> ₁₅				.022441	49089	
<i>d</i> ₁₆				-0.2316	49238	
<i>d</i> ₁₇				.025	62267	
<i>d</i> ₁₈					-0.262375	
<i>d</i> ₁₉					.02892	
<i>d</i> ₂₀					-0.294	

n = 17

<i>d</i> ₀	0.49989	65784	61135	97460	24139	91377
<i>d</i> ₁	.01449	59974	65246	53669	31342	76619
<i>d</i> ₂	-0.342	61031	30973	67782	87525	80644
<i>d</i> ₃	.0.41	26989	35462	78445	60524	58935
<i>d</i> ₄		-0.63836	27410	23481	56338	25356
<i>d</i> ₅		-0.7117	44699	41908	91673	03360
<i>d</i> ₆		-0.93	64301	73893	80901	80159
<i>d</i> ₇		.0.10	11446	21730	54924	13183
<i>d</i> ₈			-0.12364	19678	70371	23659
<i>d</i> ₉			-0.1311	73204	21503	83243
<i>d</i> ₁₀			-0.15	38252	77188	37359
<i>d</i> ₁₁				.0.161262	09253	41540
<i>d</i> ₁₂				-0.1842	12547	71133
<i>d</i> ₁₃				-0.191	42203	30276
<i>d</i> ₁₄					-0.214853	68362
<i>d</i> ₁₅					.0.22167	46192
<i>d</i> ₁₆					-0.245	83888
<i>d</i> ₁₇					.0.25	20568
<i>d</i> ₁₈						-0.27732
<i>d</i> ₁₉						.0.2826
<i>d</i> ₂₀						-0.291

n = 18

<i>d</i> ₀	0.49990	75205	19477	48462	38400	06143
<i>d</i> ₁	.01370	13610	53384	89937	61200	12457
<i>d</i> ₂	-.0 ₃ 38	03933	91259	96076	45527	65324
<i>d</i> ₃	-.0 ₄ 1	06997	97313	26394	44268	44607
<i>d</i> ₄		-.0 ₆ 3048	63628	68914	75868	53764
<i>d</i> ₅		-.0 ₈ 87	96993	12074	05317	12192
<i>d</i> ₆		-.0 ₉ 2	57021	71653	53781	28886
<i>d</i> ₇			.0 ₁₁ 7601	83147	94563	39533
<i>d</i> ₈			-.0 ₁₂ 227	55338	81348	34116
<i>d</i> ₉			.0 ₁₄ 6	89234	51704	54119
<i>d</i> ₁₀			-.0 ₁₅	21118	83108	18588
<i>d</i> ₁₁				-.0 ₁₇ 654	47218	87685
<i>d</i> ₁₂				-.0 ₁₈ 20	50828	82790
<i>d</i> ₁₃				.0 ₂₀	64965	43843
<i>d</i> ₁₄					-.0 ₂₁ 2079	92470
<i>d</i> ₁₅					.0 ₂₃ 67	28568
<i>d</i> ₁₆					-.0 ₂₄ 2	19890
<i>d</i> ₁₇						-.0 ₂₆ 7258
<i>d</i> ₁₈						-.0 ₂₇ 242
<i>d</i> ₁₉						.0 ₂₉ 8

n = 19

<i>d</i> ₀	0.49991	68139	68100	10871	35781	20729
<i>d</i> ₁	.01298	93335	66939	69397	22504	02531
<i>d</i> ₂	-0.334	16625	59780	56229	80797	83799
<i>d</i> ₃	.05	90991	25174	54684	37252	33226
<i>d</i> ₄		-0.62453	10737	94933	34978	15846
<i>d</i> ₅		.0866	93733	84520	88215	85121
<i>d</i> ₆		-0.91	84830	29942	84032	86967
<i>d</i> ₇		.0115163	52717	13211	63599	
<i>d</i> ₈		-0.12145	91574	69196	48905	
<i>d</i> ₉		.0144	17015	72530	51883	
<i>d</i> ₁₀		-0.15	12050	58313	40097	
<i>d</i> ₁₁			.017352	02921	52436	
<i>d</i> ₁₂			-0.1810	39377	06790	
<i>d</i> ₁₃			.020	31009	83492	
<i>d</i> ₁₄				-0.22934	68525	
<i>d</i> ₁₅				.02328	45630	
<i>d</i> ₁₆				-0.25	87488	
<i>d</i> ₁₇					.0262716	
<i>d</i> ₁₈					-0.2885	
<i>d</i> ₁₉					.0293	

n = 20

<i>d</i> ₀	0.49992	47740	47568	15571	94026	63730
<i>d</i> ₁	.01234	76674	46325	51533	17903	94054
<i>d</i> ₂	-0.30	85590	45564	34971	68719	19278
<i>d</i> ₃	.05	78024	03201	63418	72036	91697
<i>d</i> ₄		-0.61996	11568	78834	62091	68508
<i>d</i> ₅		.0851	65809	40859	55056	90412
<i>d</i> ₆		-0.91	35211	12319	33813	09657
<i>d</i> ₇		.0113578	75470	30275	03623	
<i>d</i> ₈		-0.1395	76775	40122	49614	
<i>d</i> ₉		.0142	59057	05423	85506	
<i>d</i> ₁₀			-0.167082	39096	35531	
<i>d</i> ₁₁			.017195	65513	56826	
<i>d</i> ₁₂			-0.195	46068	52538	
<i>d</i> ₁₃			.020	15394	46383	
<i>d</i> ₁₄				-0.22438	28846	
<i>d</i> ₁₅				.02312	59940	
<i>d</i> ₁₆				-0.25	36564	
<i>d</i> ₁₇					.0261071	
<i>d</i> ₁₈					-0.2832	
<i>d</i> ₁₉					.0291	

n = 21

<i>d</i> ₀	0.49993	16441	09513	89342	89211	09944
<i>d</i> ₁	.01176	64226	02587	64146	32927	22505
<i>d</i> ₂	-.0 ₃ 28	00432	82930	92823	43912	12050
<i>d</i> ₃	.0 ₅	67407	87998	17536	46102	95785
<i>d</i> ₄		-.0 ₆ 1640	73297	65477	77652	92006
<i>d</i> ₅		.0 ₈ 40	37771	79491	82918	71939
<i>d</i> ₆		-.0 ₉ 1	00451	23289	95382	54717
<i>d</i> ₇		.0 ₁₁ 2525	86254	47969	08672	
<i>d</i> ₈		-.0 ₁₃ 64	18519	38050	47600	
<i>d</i> ₉		.0 ₁₄ 1	64800	71879	51145	
<i>d</i> ₁₀			-.0 ₁₆ 4274	73952	79449	
<i>d</i> ₁₁			.0 ₁₇ 111	99863	26273	
<i>d</i> ₁₂			-.0 ₁₉ 2	96342	30166	
<i>d</i> ₁₃				.0 ₂₁ 7917	30059	
<i>d</i> ₁₄				-.0 ₂₂ 213	54399	
<i>d</i> ₁₅				.0 ₂₄ 5	81364	
<i>d</i> ₁₆				-.0 ₂₅	15973	
<i>d</i> ₁₇					.0 ₂₇ 443	
<i>d</i> ₁₈					-.0 ₂₈ 12	

n = 22

d_0	0.49993	76144	26027	46952	.38495	87970
d_1	-0.01123	74473	61737	38153	68775	75667
d_2	-0.325	53051	21327	19167	32267	24330
d_3	.05	58633	28868	66265	79506	33207
d_4		-0.61361	01969	06467	46394	93707
d_5		.0831	92719	28675	75913	73982
d_6		-0.10	75678	46392	23468	19859
d_7		.011812	32743	99171	35126	
d_8		-0.1343	84195	98037	17345	
d_9		.0141	07119	14499	06323	
d_{10}			-0.162643	02878	72922	
d_{11}			.01865	84566	87151	
d_{12}			-0.191	65605	63839	
d_{13}				.0214204	13854	
d_{14}				-0.22107	71193	
d_{15}				.0242	78462	
d_{16}					-0.267263	
d_{17}					.027191	
d_{18}					-0.295	

n = 23

<i>d</i> ₀	0.49994	28354	98693	55310	83704	31386
<i>d</i> ₁	.01075	39935	08228	27041	47489	15627
<i>d</i> ₂	-0 ₃ 23	37056	49498	26397	87076	93643
<i>d</i> ₃	.0 ₅	51317	69645	73222	26413	99254
<i>d</i> ₄		-0 ₆ 1138	44238	14035	10196	39414
<i>d</i> ₅		.0 ₈ 25	51213	25884	61244	06297
<i>d</i> ₆		-0 ₁₀	57745	56447	99772	23030
<i>d</i> ₇		.0 ₁₁ 1319	98694	04164	40209	
<i>d</i> ₈		-0 ₁₃ 30	46782	78996	07798	
<i>d</i> ₉		-0 ₁₅	71002	68902	48985	
<i>d</i> ₁₀			-0 ₁₆ 1670	35631	09873	
<i>d</i> ₁₁			.0 ₁₈ 39	66278	13196	
<i>d</i> ₁₂		/	-0 ₂₀	95046	22252	
<i>d</i> ₁₃				.0 ₂₁ 2298	27243	
<i>d</i> ₁₄				-0 ₂₃ 56	06862	
<i>d</i> ₁₅				.0 ₂₄ 1	37983	
<i>d</i> ₁₆					-0 ₂₆ 3425	
<i>d</i> ₁₇					.0 ₂₈ 86	
<i>d</i> ₁₈					-0 ₂₉ 2	

n = 24

<i>d</i> ₀	0.49994	74276	31398	40823	28114	58076
<i>d</i> ₁	.01031	04266	71784	28330	38064	88195
<i>d</i> ₂	-.0 ₃ 21	47355	30392	66947	57579	24014
<i>d</i> ₃	.0 ₅	45170	07062	84934	27925	41966
<i>d</i> ₄		-.0 ₇ 959	55302	42733	72780	96506
<i>d</i> ₅		.0 ₈ 20	58300	87360	69493	10395
<i>d</i> ₆		-.0 ₁₀	44577	87795	44096	04822
<i>d</i> ₇			.0 ₁₂ 974	64916	99079	82771
<i>d</i> ₈			-.0 ₁₃ 21	51009	06418	14166
<i>d</i> ₉			.0 ₁₅	47912	26801	59764
<i>d</i> ₁₀				-.0 ₁₆ 1076	97698	97931
<i>d</i> ₁₁				.0 ₁₈ 24	42671	93521
<i>d</i> ₁₂				-.0 ₂₀	55894	10211
<i>d</i> ₁₃					.0 ₂₁ 1290	18336
<i>d</i> ₁₄					-.0 ₂₃ 30	03748
<i>d</i> ₁₅					.0 ₂₅	70525
<i>d</i> ₁₆						-.0 ₂₆ 1670
<i>d</i> ₁₇						.0 ₂₈ 40
<i>d</i> ₁₈						-.0 ₂₉ 1

n = 25

<i>d</i> ₀	0.49995	14879	19156	07409	19444	17838
<i>d</i> ₁	.0 ₂ 990	20055	96695	26966	57277	24032
<i>d</i> ₂	-.0 ₃ 19	79847	00888	56154	90906	64852
<i>d</i> ₃	.0 ₅	39966	34714	66712	71310	73049
<i>d</i> ₄		-.0 ₇ 814	45519	41873	01036	46950
<i>d</i> ₅		.0 ₈ 16	75344	49356	94184	91919
<i>d</i> ₆		-.0 ₁₀	34782	29974	21437	48431
<i>d</i> ₇			.0 ₁₂ 728	75517	16801	23924
<i>d</i> ₈			-.0 ₁₃ 15	40724	88702	91789
<i>d</i> ₉			.0 ₁₅	32865	37541	89880
<i>d</i> ₁₀				-.0 ₁₇ 707	24722	85647
<i>d</i> ₁₁				.0 ₁₈ 15	35222	78622
<i>d</i> ₁₂				-.0 ₂₀	33611	40934
<i>d</i> ₁₃					.0 ₂₂ 742	10316
<i>d</i> ₁₄					-.0 ₂₃ 16	52155
<i>d</i> ₁₅					.0 ₂₅	37084
<i>d</i> ₁₆						-.0 ₂₇ 839
<i>d</i> ₁₇						.0 ₂₈ 19

n = 26

<i>d</i> ₀	0.49995	50954	20341	91191	10809	02262
<i>d</i> ₁	.0 ₂ 952	47119	06459	62982	04518	21875
<i>d</i> ₂	-.0 ₃ 18	31200	26865	13754	49163	51369
<i>d</i> ₃	.0 ₅	35532	12176	54081	79766	16381
<i>d</i> ₄		-.0 ₇ 695	77155	06122	16834	02313
<i>d</i> ₅		.0 ₈ 13	74770	43131	97410	12340
<i>d</i> ₆		-.0 ₁₀	27407	43316	36793	93052
<i>d</i> ₇		.0 ₁₂ 551	23510	55420	28766	
<i>d</i> ₈		-.0 ₁₃ 11	18382	62572	93674	
<i>d</i> ₉		.0 ₁₅	22886	68251	81454	
<i>d</i> ₁₀			-.0 ₁₇ 472	35360	07309	
<i>d</i> ₁₁			.0 ₁₉ 9	83095	97450	
<i>d</i> ₁₂			-.0 ₂₀	20631	01645	
<i>d</i> ₁₃			.0 ₂₂ 436		50831	
<i>d</i> ₁₄			-.0 ₂₄ 9		31027	
<i>d</i> ₁₅			.0 ₂₅		20016	
<i>d</i> ₁₆				-.0 ₂₇ 434		
<i>d</i> ₁₇					.0 ₂₉ 9	

n = 27

d_0	0.49995	83150	32556	55655	23951	89992
d_1	.0 ₂ 17	51173	80530	82711	50199	20540
d_2	-0 ₃ 16	98686	13130	44801	47044	27058
d_3	.0 ₅	31730	27051	63341	16424	38272
d_4		-0 ₇ 597	93703	43821	58807	30149
d_5		.0 ₈ 11	36629	83706	30345	59746
d_6		-0 ₁₀	21793	37347	89856	17424
d_7		.0 ₁₂ 421	43504	40858	84205	
d_8		-0 ₁₄ 8	21855	96523	43137	
d_9		.0 ₁₅	16161	32520	65549	
d_{10}			-0 ₁₇ 320	42840	12447	
d_{11}			.0 ₁₉ 6	40492	03743	
d_{12}			-0 ₂₀	12905	66820	
d_{13}			.0 ₂₂ 262		11151	
d_{14}				-0 ₂₄ 5	36519	
d_{15}				.0 ₂₅	11067	
d_{16}					-0 ₂₇ 230	
d_{17}					.0 ₂₉ 5	

n = 28

<i>d</i> ₀	0.49996	12004	30562	21298	60435	26779
<i>d</i> ₁	.0 ₂ 885	02794	40241	60834	68378	30822
<i>d</i> ₂	-.0 ₃ 15	80051	89929	08078	15081	58261
<i>d</i> ₃	.0 ₅	28451	97489	36554	46981	33400
<i>d</i> ₄		-.0 ₇ 516	70752	21450	12514	10905
<i>d</i> ₅		.0 ₉	46307	12931	59214	21863
<i>d</i> ₆		-.0 ₁₀	17475	80389	72101	08244
<i>d</i> ₇			-.0 ₁₂ 325	40322	70755	66100
<i>d</i> ₈			-.0 ₁₄ 6	10867	92721	85653
<i>d</i> ₉			.0 ₁₅	11560	42951	37151
<i>d</i> ₁₀				-.0 ₁₇ 220	52687	61325
<i>d</i> ₁₁				.0 ₁₉ 4	24003	17323
<i>d</i> ₁₂					-.0 ₂₁ 8215	90247
<i>d</i> ₁₃					.0 ₂₂ 160	42723
<i>d</i> ₁₄					-.0 ₂₄ 3	15642
<i>d</i> ₁₅						.0 ₂₆ 6257
<i>d</i> ₁₆						-.0 ₂₇ 125
<i>d</i> ₁₇						.0 ₂₉ 3

n = 29

<i>d</i> ₀	0.49996	37963	16978	38157	69576	10228
<i>d</i> ₁	.0 ₂ 854	76580	87697	76973	28282	82701
<i>d</i> ₂	-0 ₃ 14	73424	76558	61825	49209	67835
<i>d</i> ₃	.0 ₅	25610	13357	88319	91930	29885
<i>d</i> ₄		-0 ₇ 448	81337	89741	98250	67321
<i>d</i> ₅		.0 ₉ 7	92968	56878	49567	36972
<i>d</i> ₆		-0 ₁₀	14123	67510	01416	07123
<i>d</i> ₇			.0 ₁₂ 253	57451	24508	49079
<i>d</i> ₈			-0 ₁₄ 4	58874	01246	67584
<i>d</i> ₉				.0 ₁₆ 8369	01768	39602
<i>d</i> ₁₀				-0 ₁₇ 153	81943	83129
<i>d</i> ₁₁				.0 ₁₉ 2	84882	42188
<i>d</i> ₁₂					-0 ₂₁ 5316	17469
<i>d</i> ₁₃					.0 ₂₃ 99	94774
<i>d</i> ₁₄					-0 ₂₄ 1	89299
<i>d</i> ₁₅						.0 ₂₆ 3611
<i>d</i> ₁₆						-0 ₂₈ 69
<i>d</i> ₁₇						.0 ₂₉ 1

n = 30

<i>d</i> ₀	0.49996	61401	63493	40969	01799	09275
<i>d</i> ₁	.0 ₂ 826	50493	29711	70954	18628	17038
<i>d</i> ₂	-.0 ₃ 13	77237	45310	99787	61497	07949
<i>d</i> ₃	.0 ₅	23134	46989	09259	76344	98911
<i>d</i> ₄		-.0 ₇ 391	71195	31811	26013	83505
<i>d</i> ₅		.0 ₉ 6	68496	35746	15710	42206
<i>d</i> ₆		-.0 ₁₀	11498	04154	99744	91895
<i>d</i> ₇		.0 ₁₂ 199		30044	27998	15400
<i>d</i> ₈			-.0 ₁₄ 3	48111	60841	01533
<i>d</i> ₉				.0 ₁₆ 6126	61051	27722
<i>d</i> ₁₀				-.0 ₁₇ 108	63722	33413
<i>d</i> ₁₁				.0 ₁₉ 1	94069	91941
<i>d</i> ₁₂					-.0 ₂₁ 3492	39256
<i>d</i> ₁₃					.0 ₂₃ 63	30482
<i>d</i> ₁₄					-.0 ₂₄ 1	15575
<i>d</i> ₁₅						.0 ₂₆ 2125
<i>d</i> ₁₆						-.0 ₂₈ 39
<i>d</i> ₁₇						.0 ₂₉ 1

n = 31

<i>d</i> ₀	0.49996	82635	70250	98919	28166	19019
<i>d</i> ₁	.02800	05313	91635	05885	64851	18990
<i>d</i> ₂	-.0312	90170	29745	04663	11830	41340
<i>d</i> ₃	.05	20967	85936	12621	15044	20623
<i>d</i> ₄		-.07343	40859	23820	19892	74309
<i>d</i> ₅		.095	66746	26252	51298	83055
<i>d</i> ₆			-.0119424	45775	81022	63434
<i>d</i> ₇			.012157	90045	71076	12899
<i>d</i> ₈			-.0142	66525	15670	55434
<i>d</i> ₉				.0164531	98806	56243
<i>d</i> ₁₀				-.01877	62514	57892
<i>d</i> ₁₁				.0191	33920	20871
<i>d</i> ₁₂					-.0212326	94578
<i>d</i> ₁₃					.02340	71823
<i>d</i> ₁₄					-.025	71750
<i>d</i> ₁₅						.0261273
<i>d</i> ₁₆						-.02823

n = 32

<i>d</i> ₀	0.49997	01933	36182	88359	92436	35771
<i>d</i> ₁	.0 ₂ 775	24209	43243	62475	47254	53350
<i>d</i> ₂	-.0 ₃ 12	11105	74497	60850	74326	57445
<i>d</i> ₃	.0 ₅	19063	54553	57388	58352	53782
<i>d</i> ₄		-.0 ₇ 302	32573	85011	35044	10546
<i>d</i> ₅		.0 ₉ 4	83023	97276	26176	15098
<i>d</i> ₆			-.0 ₁₁ 7774	20557	67591	47459
<i>d</i> ₇			.0 ₁₂ 126	03993	98312	83432
<i>d</i> ₈			-.0 ₁₄ 2	05823	84236	73243
<i>d</i> ₉				.0 ₁₆ 3385	23225	47726
<i>d</i> ₁₀				-.0 ₁₈ 56	07328	82108
<i>d</i> ₁₁				.0 ₂₀	93533	53616
<i>d</i> ₁₂					-.0 ₂₁ 1571	05374
<i>d</i> ₁₃					.0 ₂₃ 26	57023
<i>d</i> ₁₄					-.0 ₂₅	45243
<i>d</i> ₁₅						.0 ₂₇ 776
<i>d</i> ₁₆						-.0 ₂₈ 13

n = 33

<i>d</i> ₀	0.49997	19523	08451	71102	68177	84255
<i>d</i> ₁	.0 ₂ 751	92372	27637	53888	73752	31908
<i>d</i> ₂	-.0 ₃ 11	39092	31534	68927	04561	29178
<i>d</i> ₃	.0 ₅	17383	00945	96041	35652	01009
<i>d</i> ₄		-.0 ₇ 267	20617	16820	46141	35342
<i>d</i> ₅		.0 ₉ 4	13711	35119	77283	74195
<i>d</i> ₆			-.0 ₁₁ 6451	35893	70925	20820
<i>d</i> ₇			.0 ₁₂ 101	31660	39118	21925
<i>d</i> ₈			-.0 ₁₄ 1	60235	22891	63319
<i>d</i> ₉				.0 ₁₆ 2551	85016	69252
<i>d</i> ₁₀				-.0 ₁₈ 40	92074	03580
<i>d</i> ₁₁				.0 ₂₀	66068	41231
<i>d</i> ₁₂					-.0 ₂₁ 1073	93220
<i>d</i> ₁₃					.0 ₂₃ 17	57368
<i>d</i> ₁₄					-.0 ₂₅	28948
<i>d</i> ₁₅						.0 ₂₇ 480
<i>d</i> ₁₆						-.0 ₂₉ 8

n = 34

<i>d</i> ₀	0.49997	35600	61604	03366	13365	71789
<i>d</i> ₁	.0 ₂ 729	96724	77055	18546	47516	22575
<i>d</i> ₂	-0 ₃ 10	73315	84370	16099	21754	25065
<i>d</i> ₃	.0 ₅	15894	32514	11254	39714	29958
<i>d</i> ₄		-0 ₇ 237	04078	64551	00150	49727
<i>d</i> ₅		.0 ₉ 3	55996	71457	14588	27780
<i>d</i> ₆		-0 ₁₁ 5383	76112	85410	75910	
<i>d</i> ₇		.0 ₁₃ 81	98169	76371	49510	
<i>d</i> ₈		-0 ₁₄ 1	25693	64847	66107	
<i>d</i> ₉		.0 ₁₆ 1940	20925	12418		
<i>d</i> ₁₀		-0 ₁₈ 30	15058	24345		
<i>d</i> ₁₁		.0 ₂₀	47165	82321		
<i>d</i> ₁₂			-0 ₂₂ 742	70443		
<i>d</i> ₁₃			.0 ₂₃ 11	77155		
<i>d</i> ₁₄			-0 ₂₅	18778		
<i>d</i> ₁₅				.0 ₂₇ 301		
<i>d</i> ₁₆					-0 ₂₉ 5	

n = 35

<i>d</i> ₀	0.49997	50334	44358	32502	31614	34201
<i>d</i> ₁	.0 ₂ 709	25673	66493	19754	66537	46034
<i>d</i> ₂	-.0 ₃ 10	13076	36919	49207	83486	17378
<i>d</i> ₃	.0 ₅	14570	87976	26657	31406	99578
<i>d</i> ₄		-.0 ₇ 211	01418	01070	27953	67113
<i>d</i> ₅		.0 ₉ 3	07678	16317	04985	64495
<i>d</i> ₆			-.0 ₁₁ 4516	65868	11197	11394
<i>d</i> ₇			.0 ₁₃ 66	74970	70950	08896
<i>d</i> ₈			-.0 ₁₅	99304	38592	38023
<i>d</i> ₉				.0 ₁₆ 1487	13202	94071
<i>d</i> ₁₀				-.0 ₁₈ 22	41647	46150
<i>d</i> ₁₁					.0 ₂₀	34009
<i>d</i> ₁₂						-.0 ₂₂ 519
<i>d</i> ₁₃						.0 ₂₄ 7
<i>d</i> ₁₄						-.0 ₂₅
<i>d</i> ₁₅						.0 ₂₇ 192
<i>d</i> ₁₆						-.0 ₂₉ 3

n = 36

<i>d</i> ₀	0.49997	63870	22708	95104	40127	80620
<i>d</i> ₁	.0 ₂ 689	68905	31844	41529	69701	83194
<i>d</i> ₂	-.0 ₄ 9	57769	43755	76882	25363	76681
<i>d</i> ₃	.0 ₅	13390	37009	83993	80384	78029
<i>d</i> ₄		-.0 ₇ 188	46330	36036	89504	31623
<i>d</i> ₅		.0 ₉ 2	67018	77863	70274	10621
<i>d</i> ₆			-.0 ₁₁ 3808	16112	20814	82475
<i>d</i> ₇			.0 ₁₃ 54	66692	07087	65183
<i>d</i> ₈			-.0 ₁₅	78985	35265	88066
<i>d</i> ₉				-.0 ₁₆ 1148	57044	56281
<i>d</i> ₁₀				-.0 ₁₈ 16	80869	26706
<i>d</i> ₁₁				.0 ₂₀	24754	33175
<i>d</i> ₁₂					-.0 ₂₂ 366	84691
<i>d</i> ₁₃					.0 ₂₄ 5	47029
<i>d</i> ₁₄						-.0 ₂₆ 8207
<i>d</i> ₁₅						.0 ₂₇ 124
<i>d</i> ₁₆						-.0 ₂₉ 2

n = 37

<i>d</i> ₀	0.49997	76334	41220	86107	89144	22884
<i>d</i> ₁	.0 ₂ 671	17213	88404	16622	64628	24455
<i>d</i> ₂	-.0 ₄ 9	06870	87987	47190	91360	53779
<i>d</i> ₃	.0 ₅	12334	00984	79687	21990	89419
<i>d</i> ₄		-.0 ₇ 168	84578	20605	09925	75408
<i>d</i> ₅		.0 ₉ 2	32639	03389	33903	92017
<i>d</i> ₆			-.0 ₁₁ 3225	97004	71187	40226
<i>d</i> ₇			.0 ₁₃ 45	01963	57751	22342
<i>d</i> ₈			-.0 ₁₅	63224	53793	85075
<i>d</i> ₉				.0 ₁₇ 893	48890	27072
<i>d</i> ₁₀				-.0 ₁₈ 12	70544	54356
<i>d</i> ₁₁				.0 ₂₀	18178	78922
<i>d</i> ₁₂					-.0 ₂₂ 261	69241
<i>d</i> ₁₃					.0 ₂₄ 3	79005
<i>d</i> ₁₄						-.0 ₂₆ 5522
<i>d</i> ₁₅						.0 ₂₈ 1
<i>d</i> ₁₆						-.0 ₂₉ 1

n = 38

<i>d</i> ₀	0.49997	87837	19331	70565	27416	90713
<i>d</i> ₁	.0 ₂ 653	62356	45496	72322	19668	24724
<i>d</i> ₂	-.0 ₄ 8	59924	34811	39067	22205	66350
<i>d</i> ₃	.0 ₅	11385	89930	24665	03643	38874
<i>d</i> ₄		-.0 ₇ 151	71545	64645	96739	58284
<i>d</i> ₅		.0 ₉ 2	03436	16251	74947	49280
<i>d</i> ₆			-.0 ₁₁ 2745	00011	60473	65723
<i>d</i> ₇			.0 ₁₃ 37	26937	72035	91977
<i>d</i> ₈			-.0 ₁₅	50913	94749	92511
<i>d</i> ₉				.0 ₁₇ 699	80209	54985
<i>d</i> ₁₀				-.0 ₁₉ 9	67712	11924
<i>d</i> ₁₁				.0 ₂₀	13462	56847
<i>d</i> ₁₂					-.0 ₂₂ 188	40764
<i>d</i> ₁₃					.0 ₂₄ 2	65238
<i>d</i> ₁₄						-.0 ₂₆ 3756
<i>d</i> ₁₅						.0 ₂₈ 53
<i>d</i> ₁₆						-.0 ₂₉ 1

n = 39

<i>d</i> ₀	0.49997	98474	95686	93118	27964	25927
<i>d</i> ₁	.0 ₂ 636	96930	36190	36148	33210	11161
<i>d</i> ₂	-.0 ₄ 8	16531	05122	56479	30746	82168
<i>d</i> ₃	.0 ₅	10532	52088	15516	93428	01926
<i>d</i> ₄		-.0 ₇ 136	70336	32595	83826	48313
<i>d</i> ₅		.0 ₉ 1	78523	24036	54863	35911
<i>d</i> ₆			-.0 ₁₁ 2345	63315	85262	26320
<i>d</i> ₇			.0 ₁₃ 31	00672	66385	89479
<i>d</i> ₈			-.0 ₁₅	41234	68072	66277
<i>d</i> ₉				.0 ₁₇ 551	64579	09735
<i>d</i> ₁₀				-.0 ₁₉ 7	42384	16768
<i>d</i> ₁₁					.0 ₂₀	10049
<i>d</i> ₁₂						-.0 ₂₂ 136
<i>d</i> ₁₃						.0 ₂₄ 1
<i>d</i> ₁₄						-.0 ₂₆ 2581
<i>d</i> ₁₅						.0 ₂₈ 36

n = 40

<i>d</i> ₀	0.49998	08332	30668	59330	22731	21769
<i>d</i> ₁	.02621	14268	76720	54402	08246	57561
<i>d</i> ₂	-0.47	76341	25875	60196	30270	82527
<i>d</i> ₃		.069762	33295	71491	60266	55660
<i>d</i> ₄		-0.7123	50284	11059	98185	07930
<i>d</i> ₅		.091	57182	80627	01495	53995
<i>d</i> ₆			-0.112012	42598	83480	29922
<i>d</i> ₇			.01325	91810	76365	04731
<i>d</i> ₈			-0.15	33576	58658	27913
<i>d</i> ₉			.017437	52287		88836
<i>d</i> ₁₀			-0.195	73425		18565
<i>d</i> ₁₁				.0217558		66432
<i>d</i> ₁₂				-0.22100		20420
<i>d</i> ₁₃					.0241	33592
<i>d</i> ₁₄					-0.261791	
<i>d</i> ₁₅						.02824

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13. ABSTRACT

The converging factor for a specific mathematical function, such as the modified Bessel function of the second kind considered in this report, is that factor by which the last term of a truncated series (usually asymptotic) approximating the function must be multiplied to compensate for the omitted terms. This converging factor for the aforementioned Bessel function is discussed herein in detail and is shown to be related to the corresponding factor for the probability integral. Tables of this factor and its reduced derivatives, correct to 30 decimal places, are included to expedite the application of this procedure to the evaluation of this Bessel function to high precision for arguments between 5 and 20, and specific examples of such applications are presented.

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